

# 1 **Internal structure of bilobate comets revealed by erosion from shear** 2 **deformation**

3  
4  
5 C. Matonti<sup>1,2\*</sup>, N. Attree<sup>1,3</sup>, O. Groussin<sup>1</sup>, L. Jorda<sup>1</sup>, S. Viseur<sup>2</sup>, S. Hviid<sup>4</sup>, S. Bouley<sup>5</sup>, D.  
6 Nebouy<sup>1</sup>, A-T. Auger<sup>1</sup>, P. Lamy<sup>1,6</sup> and, H. Sierks<sup>7</sup>, G. Naletto<sup>8,9,10</sup>, R. Rodrigo<sup>11,12</sup>, D. Koschny  
7 <sup>13</sup>, B. Davidsson<sup>14</sup>, M. A. Barucci<sup>15</sup>, J.-L. Bertaux<sup>6</sup>, I. Bertini<sup>8</sup>, D. Bodewits<sup>16</sup>, G.  
8 Cremonese<sup>17</sup>, V. Da Deppo<sup>10</sup>, S. Debei<sup>18</sup>, M. De Cecco<sup>19</sup>, J. Deller<sup>7</sup>, S. Fornasier<sup>15</sup>, M. Fulle<sup>20</sup>,  
9 P. J. Gutiérrez<sup>21</sup>, C. Güttler<sup>7</sup>, W.-H. Ip<sup>22,23</sup>, H. U. Keller<sup>24,4</sup>, L. M. Lara<sup>21</sup>, F. La Forgia<sup>8</sup>, M.  
10 Lazzarin<sup>8</sup>, A. Lucchetti<sup>17</sup>, J. J. López-Moreno<sup>21</sup>, F. Marzari<sup>8</sup>, M. Massironi<sup>25,9</sup>, S. Mottola<sup>4</sup>, N.  
11 Oklay<sup>4</sup>, M. Pajola<sup>17</sup>, L. Penasa<sup>9</sup>, F. Preusker<sup>4</sup>, H. Rickman<sup>26,27</sup>, F. Scholten<sup>4</sup>, X. Shi<sup>7</sup>, I. Toth<sup>28</sup>, C.  
12 Tubiana<sup>7</sup>, J.-B. Vincent<sup>4</sup>.

13 1 : Aix Marseille Univ, CNRS, CNES, LAM, Marseille, France, 2 : Aix-Marseille Université,  
14 CEREGE, IRD, CNRS, Ressources, Réservoirs et Hydrosystèmes/Terre et planète, Marseille,  
15 France, 3 : Faculty of natural sciences, University of Stirling, Stirling, UK, 4 : Deutsches  
16 Zentrum für Luft- und Raumfahrt (DLR), Institut für Planetenforschung, Rutherfordstraße 23,  
17 12489 Berlin, Germany. 5 : GEOPS—Géosciences Paris Sud, Université Paris-Sud, CNRS,  
18 Université Paris-Saclay, Rue du Belvédère, Bâtiment 504-509, 91405 Orsay, France. 6 :  
19 LATMOS, CNRS/UVSQ/IPSL, 11 Boulevard d’Alembert, 78280 Guyancourt, France., 7: Max  
20 Planck Institute for Solar System Research, Justus-von-Liebig-Weg 3, 37077 Göttingen,  
21 Germany. 8: University of Padova, Department of Physics and Astronomy “Galileo Galilei”, Via  
22 Marzolo 8, 35131 Padova, Italy. 9 : University of Padova, Center of Studies and Activities for  
23 Space (CISAS) “G. Colombo”, Via Venezia 15, 35131 Padova, Italy. 10 : CNR-IFN UOS Padova  
24 LUXOR, Via Trasea 7, 35131 Padova, Italy. 11 : Centro de Astrobiología, CSIC-INTA, 28850  
25 Torrejon de Ardoz, Madrid, Spain. 12: International Space Science Institute, Hallerstrasse 6,  
26 3012 Bern, Switzerland. 13 : Science Support Office, European Space Research and Technology  
27 Centre/ESA, Keplerlaan 1, Postbus 299, 2201AZ Noordwijk ZH, The Netherlands. 14 : Jet  
28 Propulsion Laboratory, M/S 183-401, 4800 Oak Grove Drive, Pasadena, CA 91109, USA. 15 :  
29 LESIA, Observatoire de Paris, Université PSL, CNRS, Univ. Paris Diderot, Sorbonne Paris Cité,  
30 Sorbonne Université, 5 Place J. Janssen, 92195 Meudon Pricipal Cedex, France. 16 : Physics  
31 Department, Auburn University, Auburn AL 36849, USA. 17 : INAF, Astronomical Observatory  
32 of Padova, Vicolo dell’Osservatorio 5, 35122 Padova, Italy. 18 : University of Padova,  
33 Department of Industrial Engineering, Via Venezia 1, 35131 Padova, Italy. 19: University of  
34 Trento, Faculty of Engineering, Via Mesiano 77, 38121 Trento, Italy. 20 : INAF Astronomical  
35 Observatory of Trieste, Via Tiepolo11, 34143 Trieste, Italy. 21 : Instituto de Astrofísica de  
36 Andalucía (CSIC), c/ Glorieta de la Astronomía s/n, 18008 Granada, Spain. 22 : Graduate  
37 Institute of Astronomy, National Central University, 300 Chung-Da Rd, Chung-Li 32054,

38 Taiwan. 23 : Space Science Institute, Macau University of Science and Technology, Avenida  
39 Wai Long, Taipa, Macau. 24 : Institut für Geophysik und extraterrestrische Physik, Technische  
40 Universität Braunschweig, Mendelssohnstr. 3, 38106 Braunschweig, Germany. 25 : University  
41 of Padova, Department of Geosciences, Via G. Gradenigo 6, 35131 Padova, Italy. 26:  
42 Department of Physics and Astronomy, Uppsala University, Box 516, 75120 Uppsala, Sweden.  
43 27 : PAS Space Research Center, Bartycka 18A, 00716 Warszawa, Poland. 28 : Konkoly  
44 Observatory, PO Box 67, 1525 Budapest, Hungary.

45

46

47 **Comets are icy bodies, eroding with time by losing gas and dust. While nucleus erosion by**  
48 **ice-sublimation has been long-known, shaping processes are still a debated question, where**  
49 **the importance of geological processes and structures remains mainly unknown. Here we**  
50 **reveal, with the example of 67P, the existence of significant mechanically-driven erosion on**  
51 **bilobate comets. It originates from a shear deformation process in the neck, possibly active**  
52 **over Gyr's, and mostly independent from the Sun distance. We report on how shear**  
53 **fracture and fault networks, characterized here for the first time, contributed to the nucleus**  
54 **mechanical erosion and how they explain 67P's strongly marked neck trough. Our 3D**  
55 **analysis proves that the nucleus interior is structured by decameter-to-hectometer shear-**  
56 **fracture networks, propagating  $\geq 500$  m below the surface, in a mechanically homogeneous**  
57 **material. This erosion process, guided by fractures, is generic and could apply to other**  
58 **bilobate comets, due to their peculiar geometry. It is a dominant process to shape the**  
59 **surface and structure the interior of bilobate comets, possibly even during their residence-**  
60 **time in the outer solar system, where water ice sublimation is negligible.**

61

62

63           Following classical dynamical scenarios<sup>1</sup>, comets were formed during the early stages of  
64 the solar system, and have been since stored far from the Sun in a very cold environment, either  
65 in the Kuiper belt or the Oort cloud. They hold clues to constrain the formation and evolution of  
66 the solar system, including insights into prebiotic molecular chemistry. Comet 67P/Churyumov-  
67 Gerasimenko (hereafter 67P), studied here, is a Jupiter family comet that originates from the  
68 Kuiper Belt.

69           During its two years orbiting comet 67P, Rosetta's cameras have acquired thousands of images  
70 revealing its bilobate nature<sup>2</sup>. Bilobate comets can be formed by, either the low-velocity (meters  
71 per second) accretion of two primordial objects<sup>3,4</sup>, or the re-aggregation of material after a later  
72 nucleus rotational breakup<sup>5</sup> or even catastrophic collision<sup>6,7</sup> that can happen multiple times. Such  
73 a configuration seems to be common for comets, since four of the seven spatially resolved nuclei  
74 are bilobate. Understanding erosion processes on cometary nuclei and how these processes  
75 modify their global shape is key to constrain their internal structure and evolution.

76           OSIRIS-NAC<sup>8</sup> image resolution (down to <20cm/px) allowed for detailed geological  
77 interpretations and in particular led to the observation of pervasive arrays of lineaments existing  
78 at all scales (from centimeters-to-hectometers), some of which have been interpreted as layers<sup>2,3</sup>  
79 and others as fractures<sup>9,10,11</sup>. While the meter-scale polygonal fractures originate from thermal  
80 stress<sup>11,12</sup>, a significant population of tens-to-hundreds of metres-scale fractures still remains from  
81 unknown origin. Lineaments are mainly observed in the Southern Hemisphere (SH), which  
82 exhibits less-to-no dust deposits compared to the Northern Hemisphere (NH)<sup>13</sup>, and therefore  
83 more continuous outcrops of brittle material, prone to fracturing (Fig. 1a).

84

85

86

## 87 **Hectometre to kilometre fracture networks near the neck**

88 In the SH neck and neck-border regions (Fig. 1a and Supplementary Figs. 1 & 3-5) we  
89 observed 2 types of tens-to-hundreds of meter long lineaments. Type-1 lineaments (green in  
90 Figure 1.) are continuous and parallel to each-other (green arrow in Fig.1c; Fig.1d, and  
91 Supplementary Fig. 5). These lineaments follow the topography contour lines and are therefore  
92 sub-parallel to the surface (Supplementary information 2.1 and Supplementary Fig. 5 and 7c-d),  
93 which agrees well with their previous interpretation and modelling as possible layers<sup>3,14</sup>.

94 Type-2 lineaments (red in Fig. 1) are composed of two sets, each with a preferential direction.  
95 They show the following attributes: (i) high interconnectivity and bent extremities (Figs. 1b2,  
96 1d2); (ii) discontinuous and curvilinear; (iii) crosscut, hence postdate, type-1 lineaments (Fig.  
97 1d1-1d2). Type-2 lineaments straightly cut across contour lines and are consequently sub-vertical  
98 to the surface (Fig 1b-c, Supplementary Figs. 3-4 and 7 and Supplementary information 1.1; 2.1  
99 and 2.2). From these characteristics, and following the geological principles of initial  
100 horizontality<sup>15</sup> and cross-cutting relationship<sup>16</sup>, type-2 lineaments cannot be primordial features  
101 such as layers, but are structural discontinuities, i.e. fractures or faults.

102

## 103 **Evidence for shear deformation**

104 In addition to the above basic attributes defining fractures, type-2 lineaments show strong  
105 evidences for shearing. Fracture terminations such as branching structures or imbricated-fans  
106 observed in the Geb and Atum regions (Fig. 1, Supplementary Figs. 3-4) are typical of fractures  
107 or faults developed in a shear context<sup>17</sup> (Figs 1b3,1d3). In the Wosret and Anhur regions at the  
108 neck's border, anastomosing fracture pattern with numerous interconnections (Fig.1b-d and  
109 Supplementary Fig. 5), along with bends of one fracture extremity toward another, indicate  
110 mechanical interaction and also suggest a shear context<sup>18,19,20</sup> (Supplementary Fig. 3b, Fig. 1b3-



111 d3). In Sobek and Neith, in the neck's central regions, we observe highly fractured sheared block  
112 structures<sup>21</sup>, alternating with unstructured areas composed of meter-scale (possibly less) blocks  
113 (Fig. 1c, 1e1-3, Supplementary information 2.3). Possible meter-scale offsets of previously  
114 formed lineaments (fractures or layers) in the Geb region clearly point to fault-like planes  
115 existing on 67P (Supplementary Fig. 3c and Supplementary information 1).

116 The evolution of these fracture and fault structures towards the neck's centre, from branching  
117 and anastomosing networks in the neck border, to sheared blocks and crushed-chaotic zones in  
118 the neck centre (Figs 1c,1e and Supplementary Fig. 6a, Supplementary information 2.3), suggests  
119 an increasing deformation gradient. This observation is fully consistent with classical  
120 fault/sheared-zone models on Earth, where maximum strain is located at its centre<sup>22</sup>  
121 (Supplementary Fig. 6b).

122 If the shearing process is a valid interpretation, fractures patterns should follow geologically  
123 significant geometry. To further assess this geometry and reinforce the evidence for shearing, we  
124 performed a quantitative analysis of lengths and directions of 2879 fractures. The fracture lengths  
125 vary from 0.5 to 450 m. The fracture cumulative length distribution follows a power law between  
126 30 and 250 m, with a power index of -2.3 (Fig. 2a). Such a distribution is typical of fractures and  
127 faults on Earth<sup>23</sup>, and this index is moreover mainly characteristic of fractures formed and/or  
128 reactivated in shear<sup>24</sup>. The 326 longest fractures (>100 m, Fig. 2c) are all strictly orientated  
129 within 35° of the neck midplane direction<sup>25</sup> (see methods) and form characteristic diamond-  
130 shaped patterns visible in both hemispheres (Supplementary Figs. 7c-d, 8c and 9) following two  
131 preferential directions separated by 30-40° (Fig. 2c-d). Such a pattern matches strikingly well  
132 with the occurrence of a Riedel-shear deformation structure between the lobes<sup>26,27,28</sup> that can exist  
133 at all scales (Supplementary Figs 8a,8d, and Supplementary information 2.4). The measured 30-  
134 40° angle between Riedel-shear fractures should correspond to the internal friction angle of the

135 material (Supplementary Fig. 8a), which fully agrees with values estimated for 67P using surface  
136 morphologies and modeling<sup>5,29,30</sup>.

137 Interestingly, the length distribution of the smallest (<30 m) fractures doesn't follow a power law,  
138 as well as the polygonal meter-scale fractures (Fig. 2b), and shows a large scattering in  
139 directions, of almost 180° (Fig. 2d), which both support a different origin for them<sup>31</sup>, likely  
140 thermal fracturing instead of "tectonic-like" shearing.

141 Finally, a stress model of 67P<sup>32</sup> has been developed (see methods), which indicates that the  
142 maximum differential stress, of up-to 450 Pa, occurs in the neck regions (Supplementary  
143 Fig. 10a). This value exceeds the estimated bulk nucleus (tensile or shear) strength of typically 1-  
144 100 Pa<sup>33,34</sup>, thus allowing fracturing. Shear stress is also maximum near the Neck centre  
145 (>100 Pa, Supplementary Fig. 10b) and in the neck's perpendicular direction, which is  
146 compatible with the location and directions of the observed shear-deformations. Such a stress is  
147 caused by torque at the neck boundary, due to the fact that the neck, plus head-lobe, are  
148 cantilevered over 67P's centre of gravity and falling onto it with a twisting motion.

149 To summarise, all the above observations and models cannot be explained by thermal  
150 processes and demonstrate the occurrence of a global shear deformation happening all around the  
151 neck, which is mostly independent of solar insolation and has been active far from the sun,  
152 possibly over Gyr's<sup>4</sup>, since 67P became bilobate.

153

#### 154 **Constraining the nucleus internal structure**

155 The global shear stress not only implies surface deformation, but also a strain in the whole  
156 nucleus interior. In order to assess this hypothesis, we studied fractures along the vertical  
157 direction, relative to the local gravity vector (i.e. along their height instead of their strike), hence  
158 probing the nucleus internal structure.

159 Fractures observed vertically on cliffs in the NH equatorial region (thereafter Bakhu,  
160 Supplementary Figs. 7a-b and 9b and Supplementary information 2.4 & 3) show maximum  
161 heights of 120-190 m. Considering a maximum length of 450 m, this gives fracture  
162 Length/Height ratios around 2.36-3.75 (Fig. 3) typical for fractures and faults on Earth, especially  
163 non-layer-restricted ones ( $1.8 < L/H < 3.8$  <sup>35</sup>).

164 Fractures on the neck borders (Wosret and Anhur) and neck deepest point (Sobek) exhibit  
165 similar patterns and sheared block structures (Figs 1c and 1e2 and Supplementary information 2.2  
166 & 2.3). This observation proves that they propagate towards the nucleus interior, being part of a  
167 same network, and proves that shearing occurs inside the nucleus, down to at least several  
168 hundreds-of-meters, i.e. the maximum neck depth (Fig. 3).

169 In both hemispheres, the neck borders exhibit cliff faces (Supplementary Figs 7c-d, 9b) that  
170 mainly follow the 2 preferential directions of fractures (Supplementary Fig. 8b). These particular  
171 cliffs are thus the remains of fractures walls, where the opposite side has been eroded. Therefore,  
172 the nucleus material breakdown, and then erosion, in the neck (i.e. neck's trough shaping) has  
173 been, at least partially, caused by shear deformation. Indeed, mechanical breakdown may act or  
174 have acted as an amplification/facilitating process for increasing sublimation, by exposing more  
175 pristine, non-dust-covered material, more prone to sublimation and may also have allowed block  
176 removal and transport/escape.

177 These observations imply that finally, from the neck border to the neck centre/bottom, we  
178 therefore observe the same shear structure over hundreds-of-metres depth, at different evolution  
179 levels, driven by mechanical erosion along underlying and pre-existing fractures (Fig.3). It goes  
180 from fractured with little mechanical erosion in the neck borders (1 in Fig. 3b-c), to partially  
181 eroded in Bakhu (2 in Fig. 3b-c), and finally to highly sheared/crushed and eroded, forming  
182 flattened areas, in the neck centre (3 in Fig. 3b-c). These 3D observations necessarily imply that:

183 (i) the nucleus interior is structured by decametre-to-hectometre fracture networks, (ii) the  
184 nucleus material remains sufficiently brittle below the surface to allow fracturing, even at  
185 several-hundreds-of-meters depths, (iii) although the nucleus exhibits layering, it is mechanically  
186 homogenous enough for fractures to propagate freely, without being stopped or damped by  
187 mechanical boundaries, i.e. that layers do not show sharp mechanical contrasts.

188

### 189 **Implications for the evolution of bilobate comets**

190 With these results, we can now propose a chronology explaining the erosion and shape  
191 evolution of 67P (Fig.4).

192 Step 1 – Following <sup>4</sup>, 67P acquired its bilobate shape roughly 4.5 Gyrs ago, in the primordial or  
193 scattered disk, with the low-velocity accretion of two cometesimals<sup>4,5</sup> (Fig. 4a). Alternatively, the  
194 bilobate shape could also originate from a more recent catastrophic collision and re-accumulation  
195 event <sup>6</sup>, although the probability of such an event drops to < 5% after 3.5 Gyr BP<sup>36</sup>, when the  
196 Kuiper belt acquired its current object density<sup>37</sup> (Fig. 4a).

197 Step 2 – Then, due to torque-stress at the lobes' boundary, which originates from an initial  
198 asymmetry of the nucleus, shear deformation starts all around 67P, leading to pervasive  
199 fracturing (i.e. type 2 lineaments, Fig. 4b). This stress originates from the geometry of the  
200 nucleus, but we cannot exclude that several later close encounters with giant planets<sup>1</sup> provided  
201 additional tidal stresses. Continuing shear deformation, possibly over Gyr's, progressively  
202 increases the fracturing level (length, connections...), hence producing even more  
203 broken/damaged material in the neck.

204 Step 3 – Hundred-thousands-to-millions of years ago<sup>1</sup>, 67P entered the giant planet region. Its  
205 temperature slowly increases and sublimation of the most volatile ices (CO, N<sub>2</sub> and CO<sub>2</sub>) starts  
206 (Fig. 4c). The still-ongoing mechanical breakdown induced by shearing continues to weaken the

207 nucleus material, setting the conditions for differential erosion focused on the lobe boundary,  
208 increasing its depth, and forming a deep “neck” (Fig 4c).

209 Along with broad volatile sublimation, outbursts induced by fracturing<sup>38</sup> (or fracture  
210 reactivation), driven by sudden confinement/pressure loss, could have also contributed to ejecting  
211 and eroding the fractured loose blocks from the nucleus, enhancing the preferential neck erosion.  
212 This process is especially plausible and most efficient for smaller blocks in the crushed-material  
213 of the neck centre. This leads to increasing cliff heights surrounding the neck, therefore leads to  
214 more probability of cliff collapse<sup>39</sup> and block fragmentation, exposing ice rich materials to the  
215 surface, which amplifies even more the preferential erosion in the neck.

216 Step 4 – Ultimately, 67P reaches the inner solar system and its Jupiter-family comet orbit, with a  
217 perihelion distance,  $q$ , inside 5 AU. At this distance, the sublimation of water starts, leading to  
218 broad erosion by sublimation, which becomes the dominant process here (Fig. 4d). Reaching its  
219 current orbit ( $q=1.2$  AU), the erosion is typically 0.4-1 m per orbit<sup>40,41</sup>. This significant erosion  
220 primarily affects the most insolated areas at perihelion, i.e. the SH<sup>40</sup>, and more precisely the lobes  
221 rather than the neck, where projected shadows limit insolation. It is responsible for the large  
222 depth difference between the SH ( $\approx 450$  m) and NH ( $\approx 930$  m), by flattening the SH neck’s flanks,  
223 giving 67P its current North-South asymmetric shape (Fig 4e).

224 The above scenario and conclusions on the internal structure are not restricted to 67P and can  
225 apply to other bilobate comets, which is likely a common shape among cometary nuclei, and  
226 could explain previous (even non-directly observed) nucleus splitting<sup>42</sup>. As shown by the recent  
227 New Horizon mission, active geological processes exist in the Kuiper belt, on long time scale,  
228 and comets are no exception. Finally, this work also brings new perspectives on the comet  
229 activity phenomenon, where deep propagating fracture and fault growth could trigger outbursts<sup>38</sup>,  
230 even at large ( $>5$  a.u.) heliocentric distances<sup>43</sup>.

231 **References:**

- 232 1 Duncan, M., Levison, H. & Dones, L. Dynamical evolution of ecliptic comets. *Comets II*,  
233 193-204 (2004).
- 234 2 Sunshine, J. M., Thomas, N., El-Maarry, M. R. & Farnham, T. L. Evidence for geologic  
235 processes on comets. *Journal of Geophysical Research-Planets* 121, 2194-2210,  
236 doi:10.1002/2016je005119 (2016).
- 237 3 Massironi, M. et al. Two independent and primitive envelopes of the bilobate nucleus of  
238 comet 67P. *Nature* 526, 402-+, doi:10.1038/nature15511 (2015).
- 239 4 Davidsson, B. J. R. et al. The primordial nucleus of comet 67P/Churyumov-Gerasimenko.  
240 *Astronomy & Astrophysics* 592, doi:10.1051/0004-6361/201526968 (2016).
- 241 5 Hirabayashi, M. et al. Fission and reconfiguration of bilobate comets as revealed by  
242 67P/Churyumov-Gerasimenko. *Nature* 534, 352-+, doi:10.1038/nature17670 (2016).
- 243 6 Schwartz, S. R. *et al.* Catastrophic disruptions as the origin of bilobate comets. *Nature*  
244 *Astronomy* 2, 379-382, doi:10.1038/s41550-018-0395-2 (2018).7 Jutzi, M. & Benz, W.  
245 Formation of bi-lobed shapes by sub-catastrophic collisions A late origin of comet 67P's  
246 structure. *Astronomy & Astrophysics* 597, doi:10.1051/0004-6361/201628964 (2017).
- 247 7 Jutzi, M. & Benz, W. Formation of bi-lobed shapes by sub-catastrophic collisions A late  
248 origin of comet 67P's structure. *Astronomy & Astrophysics* 597, doi:10.1051/0004-  
249 6361/201628964 (2017).
- 250 8 Keller, H. U. et al. OSIRIS – The Scientific Camera System Onboard Rosetta. *Space*  
251 *Science Reviews* 128, 433-506, doi:10.1007/s11214-006-9128-4 (2007).

252 9 El-Maarry, M. R. et al. Fractures on comet 67P/Churyumov-Gerasimenko observed by  
253 Rosetta/OSIRIS. *Geophysical Research Letters* 42, 5170-5178, doi:10.1002/2015gl064500  
254 (2015).

255 10 Thomas, N. et al. The morphological diversity of comet 67P/Churyumov-Gerasimenko.  
256 *Science* 347, doi:10.1126/science.aaa0440 (2015).

257 11 Auger, A. T. et al. Meter-scale thermal contraction crack polygons on the nucleus of  
258 comet 67P/Churyumov-Gerasimenko. *Icarus*, doi:https://doi.org/10.1016/j.icarus.2017.09.037  
259 (2017).

260 12 Attree, N. *et al.* Thermal fracturing on comets. *A&A* **610** (2018).  
261

262 13 Lee, J. C. et al. Geomorphological mapping of comet 67P/Churyumov-Gerasimenko's  
263 Southern hemisphere. *Monthly Notices of the Royal Astronomical Society* 462, S573-S592,  
264 doi:10.1093/mnras/stx450 (2016).

265 14 Penasa, L. a. M. M. a. N. G. a. S. E. a. F. S. a. P. M. a. L. A. a. P. F. a. S. F. a. A three  
266 dimensional modelling of the layered structure of comet 67P/Churyumov-Gerasimenko. *Monthly*  
267 *Notices of the Royal Astronomical Society*, stx2899, doi:10.1093/mnras/stx2899 (2017).

268 15 Lyell, C. & Deshayes, G. P. *Principles of Geology: Being an Attempt to Explain the*  
269 *Former Changes of the Earth's Surface, by Reference to Causes Now in Operation.* (J. Murray,  
270 1830).

271 16 Steno, N. *De Solido Intra Solidum Naturaliter Contento.*, Vol. 78 (1669).

272 17 Twiss, R. J. & Moores, E. M. *Structural geology.* (Macmillan, 1992).

273 18 Rao, G. et al. Co-seismic Riedel shear structures produced by the 2010 M-w 6.9 Yushu  
274 earthquake, central Tibetan Plateau, China. *Tectonophysics* 507, 86-94,  
275 doi:10.1016/j.tecto.2011.05.011 (2011).

276 19 Matonti, C., Lamarche, J., Guglielmi, Y. & Marié, L. Structural and petrophysical  
277 characterization of mixed conduit/seal fault zones in carbonates: Example from the Castellás fault  
278 (SE France). *Journal of Structural Geology* 39, 103-121, doi:10.1016/j.jsg.2012.03.003 (2012).

279 20 Peacock, D. C. P., Nixon, C. W., Rotevatn, A., Sanderson, D. J. & Zuluaga, L. F.  
280 Glossary of fault and other fracture networks. *Journal of Structural Geology* 92, 12-29,  
281 doi:10.1016/j.jsg.2016.09.008 (2016).

282 21 Mukherjee, S. *Atlas of Structural Geology*. (Elsevier Science, 2015).

283 22 Mitchell, T. M. & Faulkner, D. R. The nature and origin of off-fault damage surrounding  
284 strike-slip fault zones with a wide range of displacements: A field study from the Atacama fault  
285 system, northern Chile. *Journal of Structural Geology* 31, 802-816, doi:10.1016/j.jsg.2009.05.002  
286 (2009).

287 23 Soliva, R. and Schultz, R. A. Distributed and localized faulting in extensional settings:  
288 Insight from the North Ethiopian Rift–Afar transition area. *Tectonics* 27,  
289 doi:10.1029/2007TC002148.

290 24 Cladouhos, T. T. & Marrett, R. Are fault growth and linkage models consistent with  
291 power-law distributions of fault lengths? *Journal of Structural Geology* 18, 281-293,  
292 doi:10.1016/s0191-8141(96)80050-2 (1996).

293 25 Preusker, F. et al. The global meter-level shape model of comet 67P/Churyumov-  
294 Gerasimenko. *Astronomy & Astrophysics*, 607, L1, doi:10.1051/0004-6361/201731798 (2017)

295 26 Riedel, W. Zur Mechanik Geologischer Brucherscheinungen *Central bl. F. Min. Geol.*  
296 *Und Pal* 8, 354-368 (1929).

297 27 Bartlett, W. L., Friedman, M. & Logan, J. M. Experimental folding and faulting of rocks  
298 under confining pressure .9. wrench faults in limestone layers. *Tectonophysics* 79, 255-277,  
299 doi:10.1016/0040-1951(81)90116-5 (1981).



300 28 Ahlgren, S. G. The nucleation and evolution of Riedel shear zones as deformation bands  
301 in porous sandstone. *Journal of Structural Geology* 23, 1203-1214, doi:10.1016/s0191-  
302 8141(00)00183-8 (2001).

303 29 Groussin, O. et al. Gravitational slopes, geomorphology, and material strengths of the  
304 nucleus of comet 67P/Churyumov-Gerasimenko from OSIRIS observations. *Astronomy &*  
305 *Astrophysics* 583, doi:10.1051/0004-6361/201526379 (2015).

306 30 Vincent, J. B. et al. Constraints on cometary surface evolution derived from a statistical  
307 analysis of 67P's topography. *Monthly Notices of the Royal Astronomical Society* 469, S329-  
308 S338, doi:10.1093/mnras/stx1691 (2017).

309 31 Hatton, C. G., Main, I. G. & Meredith, P. G. Non-universal scaling of fracture length and  
310 opening displacement. *Nature* 367, 160-162 (1994).

311 32 Hviid, S., Hüttig, C., Groussin, O., Mottola, S., Keller, H. U., et al. A Creaking and  
312 Cracking Comet. AAS/Division for Planetary Sciences Meeting Abstracts, Vol. 48,  
313 AAS/Division for Planetary Sciences Meeting Abstracts, 211.05 (2016).

314 33 Attree, N. *et al.* Tensile strength of 67P/Churyumov–Gerasimenko nucleus material from  
315 overhangs. *A&A* **611** (2018).

316 34 Basilevsky, A. T. et al. Estimating the strength of the nucleus material of comet 67P  
317 Churyumov–Gerasimenko. *Solar System Research* 50, 225-234,  
318 doi:10.1134/s0038094616040018 (2016).

319 35 Hooker, J. N., Laubach, S. E. & Marrett, R. Fracture-aperture size—frequency, spatial  
320 distribution, and growth processes in strata-bounded and non-strata-bounded fractures, Cambrian  
321 Mesón Group, NW Argentina. *Journal of Structural Geology* 54, 54-71,  
322 doi:https://doi.org/10.1016/j.jsg.2013.06.011 (2013).

323 36 Durda, D. D. & Stern, S. A. Collision Rates in the Present-Day Kuiper Belt and Centaur  
324 Regions: Applications to Surface Activation and Modification on Comets, Kuiper Belt Objects,

325 Centaurs, and Pluto–Charon. *Icarus* 145, 220-229, doi:<https://doi.org/10.1006/icar.1999.6333>  
326 (2000).<sup>37</sup> Weissman, P. R. & Levison, H. F. in *Pluto and Charon* 559 (1997).<sup>38</sup> Skorov,  
327 Y. V., Rezac, L., Hartogh, P., Bazilevsky, A. T. & Keller, H. U. A model of short-lived outbursts  
328 on the 67P from fractured terrains. *A&A* 593, A76 (2016).

329 39 Pajola, M. et al. The pristine interior of comet 67P revealed by the combined Aswan  
330 outburst and cliff collapse. *Nature Astronomy* 1, doi:10.1038/s41550-017-0092 (2017).

331 40 Lai, I. L. et al. Gas outflow and dust transport of comet 67P/Churyumov-Gerasimenko.  
332 *Monthly Notices of the Royal Astronomical Society* 462, S533-S546, doi:10.1093/mnras/stx332  
333 (2016).

334 41 Keller, H. U. et al. Insolation, erosion, and morphology of comet 67P/Churyumov-  
335 Gerasimenko. *A&A* 583, A34 (2015).

336 42 Boehnhardt, H. in *Comets II* (ed M.C. and Keller Festou, H.U. and Weaver, H.A.) 301-  
337 316 (2004).

338 43 Sekanina, Z., Larson, S. M., Hainaut, O., Smette, A. & West, R. M. Major outburst of  
339 periodic Comet Halley at a heliocentric distance of 14 AU. *Astronomy and Astrophysics* 263,  
340 367-386 (1992).

341  
342  
343  
344  
345  
346  
347  
348

349 **Acknowledgments**

350 OSIRIS was built by a consortium of the Max-Planck-Institut für Sonnensystemforschung,  
351 Göttingen, Germany; the CISAS University of Padova, Italy; the Laboratoire d'Astrophysique de  
352 Marseille, France; the Instituto de Astrofísica de Andalucía, CSIC, Granada, Spain; the Research  
353 and Scientific Support Department of the ESA, Noordwijk, Netherlands; the Instituto Nacional de  
354 Técnica Aeroespacial, Madrid, Spain; the Universidad Politécnica de Madrid, Spain; the  
355 Department of Physics and Astronomy of Uppsala University, Sweden; and the Institut für  
356 Datentechnik und Kommunikationsnetze der Technischen Universität Braunschweig, Germany.  
357 The support of the national funding agencies of Germany (DLR), France (CNES), Italy (ASI),  
358 Spain (MEC), Sweden (SNSB), and the ESA Technical Directorate is gratefully acknowledged.  
359 We thank the Rosetta Science Operations Centre and the Rosetta Mission Operations Centre for  
360 the successful rendezvous with comet 67P/Churyumov–Gerasimenko.

361 The authors would also like to thank the ParadigmGeo company for their support in providing the  
362 gOcad software, V. A. La Bruna for the interesting Structural Geology discussions related to this  
363 study, and Y. Guglielmi for his advices on the submission.

364

365 **Data Availability**

366 All the images analysed during the current study are available in the ESA-PSA repository  
367 (<https://archives.esac.esa.int/psa>). The data that support the findings of this study are in the supplementary  
368 information section and available from the corresponding author upon reasonable request (C.Matonti;  
369 [matonti@cerege.fr](mailto:matonti@cerege.fr)).

370

371

372 **Author contributions**

373 C.M led this study, mapped the lineaments, performed geological interpretation and wrote most  
374 of the manuscript. N.A performed the 3D projection of the lineaments as well as the statistical  
375 calculations and interpretations, and participated to the manuscript writing. O.G contributed  
376 significantly to the interpretations and to the manuscript writing. L.J provided local and global  
377 3D models and developed tool for images selection and data projection. S.V contributed to the  
378 3D statistical analysis and data importing to the Gocad software. S.H provided the 3D stress  
379 model for 67P. S.B contributed to improved design of the study, interpretations and manuscript.  
380 D.N contributed to the local and global 3D shape model creation. A-T.A contributed to the image  
381 selection and geological interpretation. P.L provided Stereo anaglyph images used for  
382 interpretation.

383 H.S., C.B., P.L., R.R., D.K. and H.R. are the lead scientists of the OSIRIS project. The other  
384 authors are all co-investigators who built and ran this instrument and made the observations  
385 possible, and associates and assistants who participated in the study.

386

387 **Figures Captions:**

388 Figure 1: Fracture pattern on 67P's SH showing fracture interpretations and comparison with  
389 typical Earth analogues/equivalents. a. NAC\_2016-01-  
390 28T05.33.00.986Z\_ID30\_1397549000\_F22 image, showing neck borders and centre regions. b.  
391 NAC\_2016-01-30T08.28.39.721Z\_ID30\_1397549200\_F22, Wosret neck border (b1) with  
392 interpreted layers (green) and digitalized fracture lineaments (red), showing anastomosing and  
393 highly interconnected pattern (b2). Earth example of a fault-zone showing similar anastomosing  
394 pattern and following typical Riedel-shear structure (with R&P planes)<sup>18</sup> (b3). c. NAC\_2016-01-

395 27T18.20.08.974Z\_ID30\_1397549000\_F22 cropped image, showing close-up of the neck centre  
396 exhibiting ridges and chaotic zones. d. Zoom on the neck border (d1) with fractures (red)  
397 crosscutting possible layers (green), highlighting the occurrence of a dense, oblique fracture set  
398 located in-between longer fractures (d2). Example of anastomosing shear fault-zone with  
399 sinistral-slip motion, including oblique shear-fractures between minor fault and sheared block<sup>19</sup>  
400 (d3). e. Close-up on brittle material ridges in the neck bottom (e1), affected by oblique fractures  
401 (in red) and chaotic/unstructured crushed zones (in grey) (e2). Image of a sheared rock-block,  
402 pinched between two minor faults, exhibiting oblique Riedel-shear fractures (e3, from<sup>21</sup>).

403  
404 Figure 2: Fracture length distribution and directions statistics. a. Cumulative length distribution  
405 plot of all the digitalized fracture. The distribution follows a power law between 30-40 and 250 m  
406 (see methods). b Cumulative length distribution of meter scale polygonal fractures (data from<sup>11</sup>)  
407 showing no evidence of a power law distribution, but exponential distribution, contrarily to  
408 tectonic shear (linked or reactivated) fractures and faults which classically exhibit power law  
409 distribution. c. Polar plot of the longest fractures ( $L > 100$  m) average directions compared to the  
410 neck middle plane ( $\theta$  angle =  $0^\circ$  means parallel to the neck). It shows scattering mainly in a  $35^\circ$   
411 range. d. Polar plot of the shorter fracture ( $L < 100$  m) directions compared to the neck middle  
412 plane. It shows large scattering over a  $> 100^\circ$  range.

413  
414 Figure 3: Block diagram revealing 67P's fractured internal structure and its evolution through  
415 increasing mechanical erosion. a. Location of the 3 views on 67P's nucleus. b. NAC images  
416 illustrating increasing erosion level along a unique fractured structure. c. Block diagrams.  
417 1: initial, non (mechanically) eroded, topographic surface, as observed in Wosret. 2: partially

418 eroded topographic surface, cut along the pre-existing fracture directions and dip angles, as  
419 observed in Hapi equatorial area (Bakhu). 3. Highly flattened surface topography, eroded along  
420 the same fracture network with increased deformation in the Neck's centre, forming lenticular  
421 shaped ridge and crushed chaotic zones, as observed in Anhur/Sobek regions.

422  
423 Figure 4: Chronology of the evolution of the shape of 67P (from primordial or collisional later  
424 event), showing the effects of the two complementary erosion processes (mechanical erosion and  
425 sublimation erosion). It highlights the contrast between the shear deformation, acting in the neck  
426 over long time scales (Gy's), and the sublimation erosion, acting on the broad nucleus over  
427 shorter time scales (My's). Double red-arrows are the symbol for shear deformation, illustrating  
428 torque at the neck, and do not imply a sense of rotation.

429

430

## 431 **Methods**

432 10 OSIRIS-NAC images, acquired between 8.3 and 70 km from the nucleus centre of  
433 mass, were used for this work in order to digitalise 2879 lineaments. Digitalisation was  
434 performed using the vector-based Adobe illustrator drawing software. Image resolution and size  
435 at the nucleus surface range respectively from 0.33 -1.23 m/px and 0.69 - 2.79 km. Digitalised  
436 lineaments are polyline objects made only of straight line combinations (Supplementary Fig. 2a).  
437 Lineaments were exported in .svg format and projected onto the SPG-SHAP7<sup>25</sup> nucleus model,  
438 using the known geometry of OSIRIS images<sup>44</sup> (Supplementary Fig. 2b). A neck axis plane was  
439 defined, using the midplane of the neck border coordinates from<sup>25</sup>, and the average distance and  
440 direction (weighted by segment lengths) of each lineament were computed relative to this.

441 The fracture cumulative length distribution fit was performed following classic  
442 recommendations from<sup>45</sup>. Our data set is composed of more than 200 measurements and ranges  
443 over 2 orders of magnitude. Nevertheless, for power law exponent determination, we only  
444 sampled fractures inside 0.5 to 25% of the image actual size, in order to avoid typical issues such  
445 as: (i) truncation effect, due to image resolution limits; (ii) length bias or censoring effect, due to  
446 the size of images/sampling area compared to the size of longest fractures; and (iii) statistical  
447 effect due to undersampling of the largest objects.

448 Cliff directions have been computed using the Gocad software (Paradigmgeo), by  
449 drawing lines parallel to the cliffs directly onto the shape model mesh triangles. Parallel-  
450 view/orthographic view was used in order to avoid perspective/parallax effect bias. Anaglyph 3D  
451 view mode has been used to better estimate depth, in order to accurately draw the line on the  
452 cliffs, minimizing error in cliffs directions.

453 The full stress tensor for 67P was computed taking into account gravity and rotational  
454 forces, considering a Young's modulus value of 50 MPa and a Poisson ratio of 0.32<sup>32</sup>. The stress  
455 model was computed using a finite element mesh composed of 2 million cells, and principal  
456 stress values and directions were mapped onto the cg-dlr\_spg-shap7-v1.0 model<sup>25</sup>.

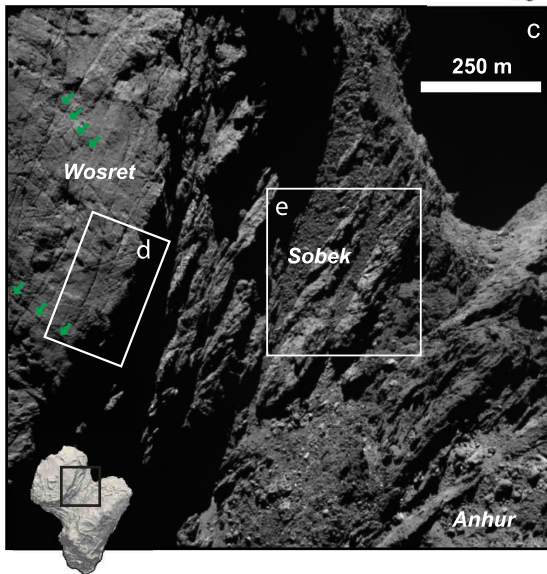
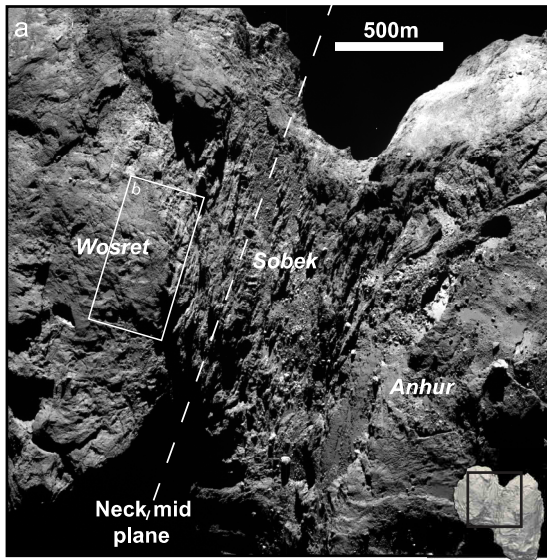
457

#### 458 **References only in Methods:**

459 44 Gaskell, R. W. a. B.-J. H. A. O. S. a. S. D. J. a. K. A. S. a. M. T. a. A. B. Characterizing  
460 and navigating small bodies with imaging data. *Meteoritics & Planetary Science* 43, 1049--1061,  
461 doi:10.1111/j.1945-5100.2008.tb00692.x (2008).

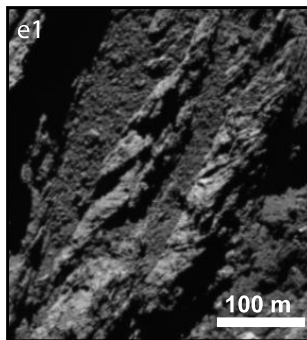
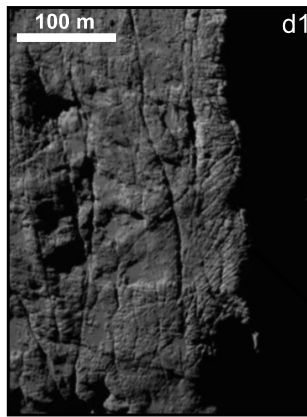
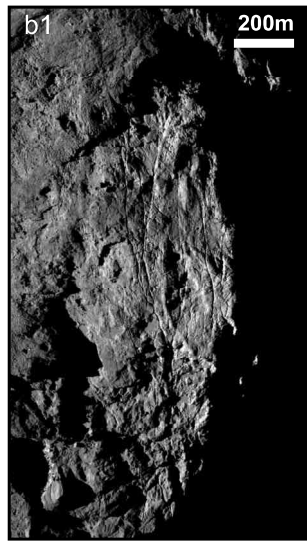
462 45 Bonnet, E. a. B. O. a. O. N. E. a. D. P. a. M. I. a. C. P. a. B. B. Scaling of fracture systems  
463 in geological media. *Reviews of Geophysics* 39, 347--383, doi:10.1029/1999RG000074 (2001).

Full images

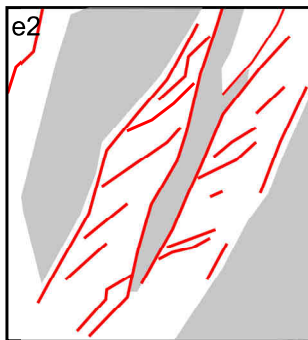
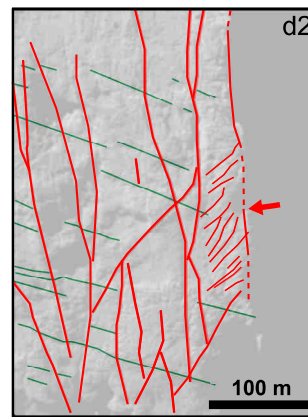
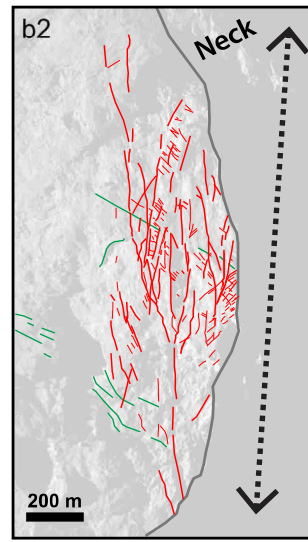


- Possible Layers
- Fractures
- Interpreted Layers lineament location
- Chaotic zones
- Nucleus topography
- Shear motion direction

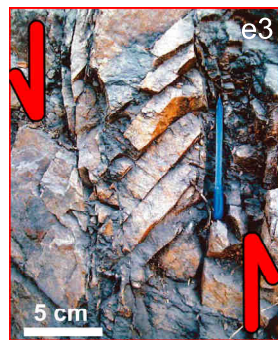
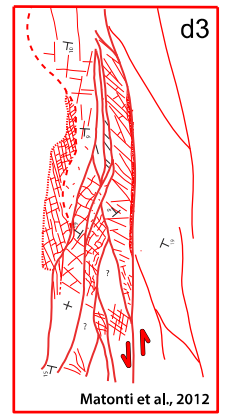
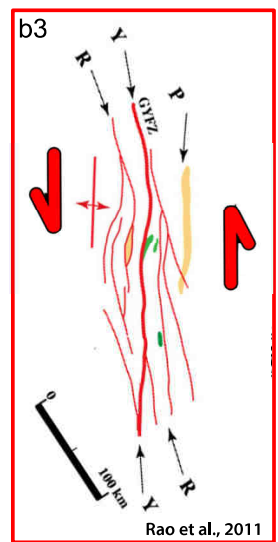
Inset images



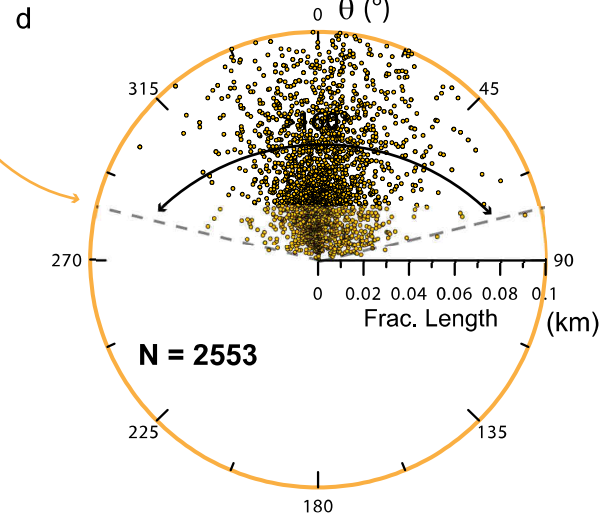
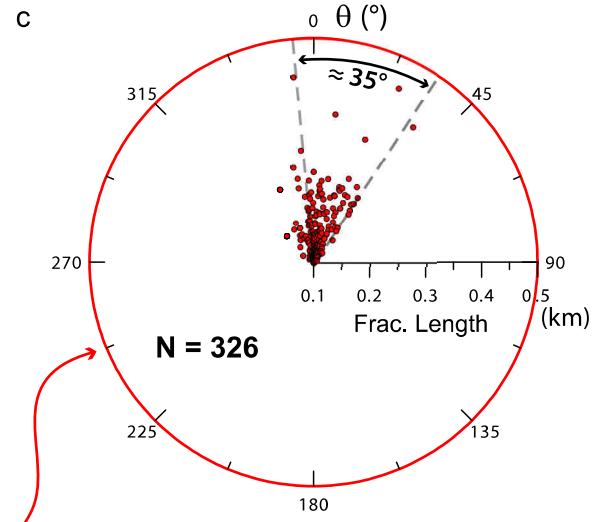
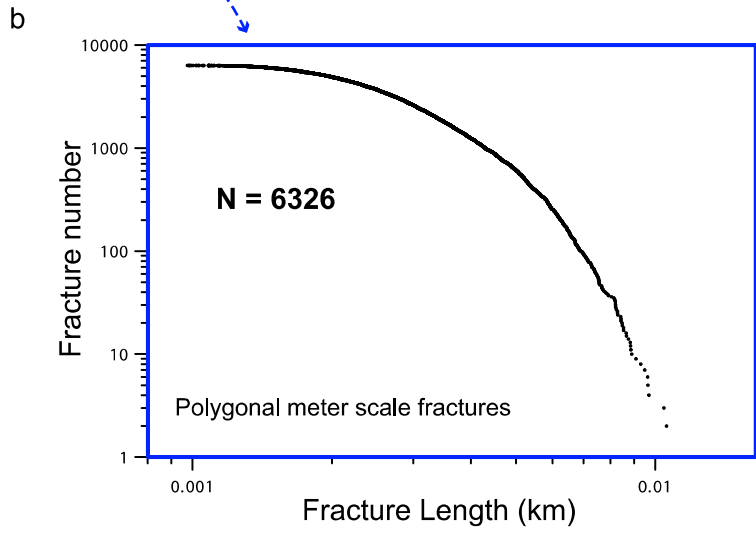
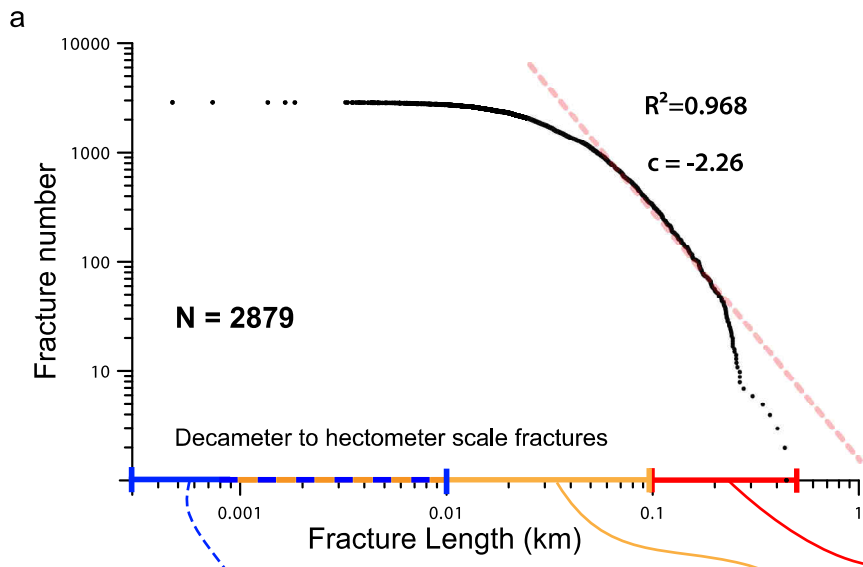
Pattern



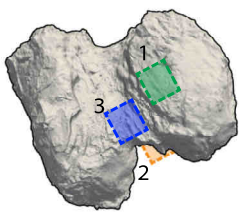
Earth analogues





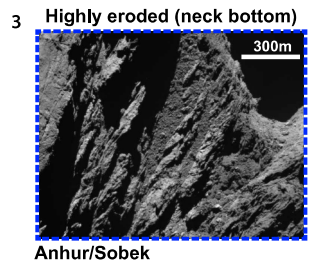
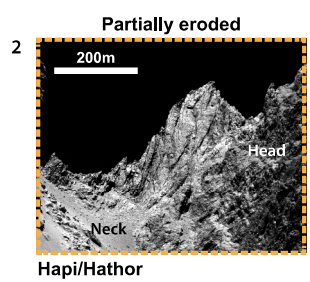
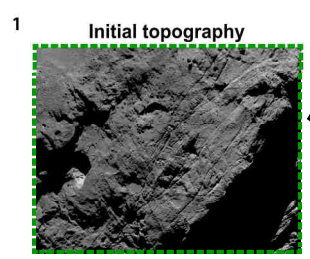


a

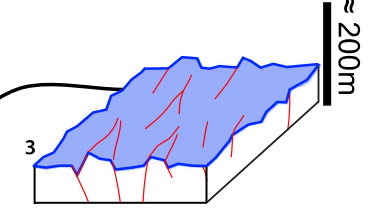
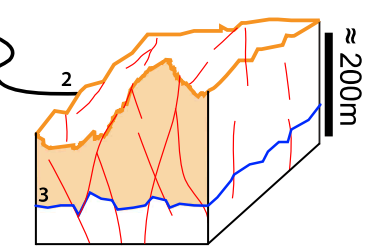
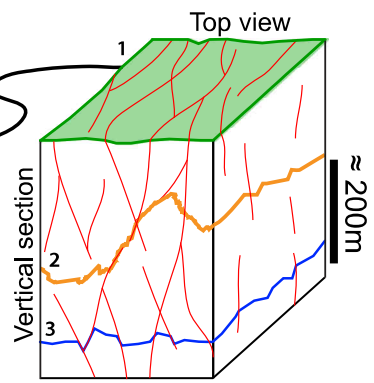


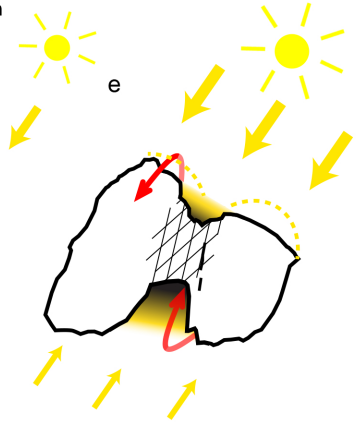
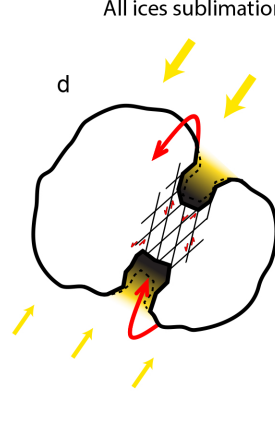
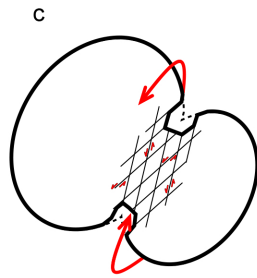
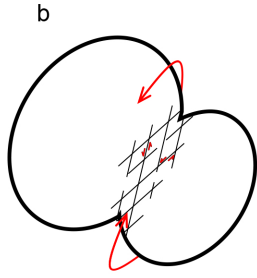
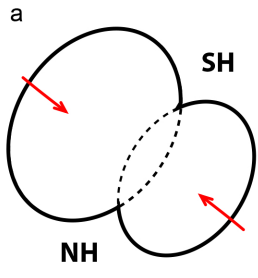
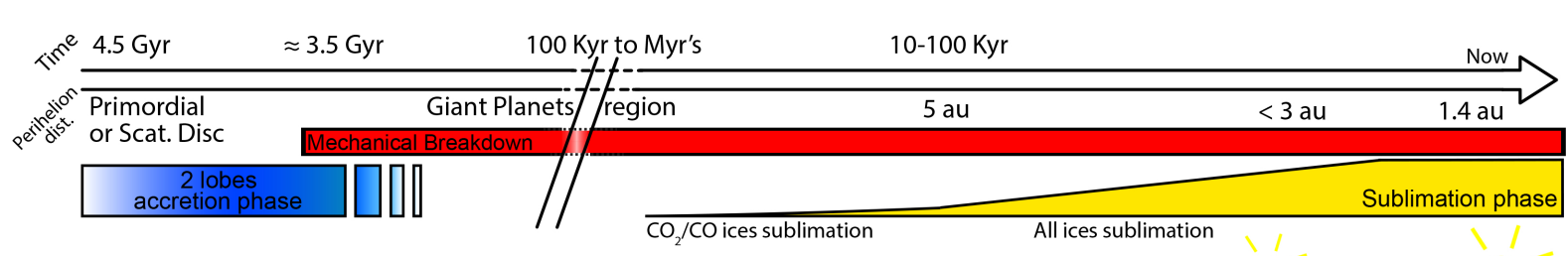
- View locations on the nucleus
- Fractures Lineaments
- Nucleus current topography surfaces

b



c





Mechanical erosion  $\gg$  Sublimation erosion

Sublimation erosion  $\gg$  Mechanical erosion